# **ETpathfinder**

#### Stefan Hild

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Maastricht University and NIKHEF (stefan.hild@maastrichtuniversity.nl) for the ETpathfinder Team

Nikhef, Maastricht University, University of Antwerp, Ghent University, Katholieke Universiteit Leuven, Université Catholique de Louvain, Hasselt University, Vrije Universiteit Brussel, Fraunhofer Institute for Laser Technology, RWTH Aachen, University of Twente, Eindhoven University of Technology, Liege Université, VITO, TNO + LAPP Annecy, Vrije Universiteit Amsterdam, Albert Einstein Institute Hannover, Birmingham University, Karlsruhe Institute of Technology (KIT) ...

> Maastricht University

#### www.etpathfinder.eu





CLAPP

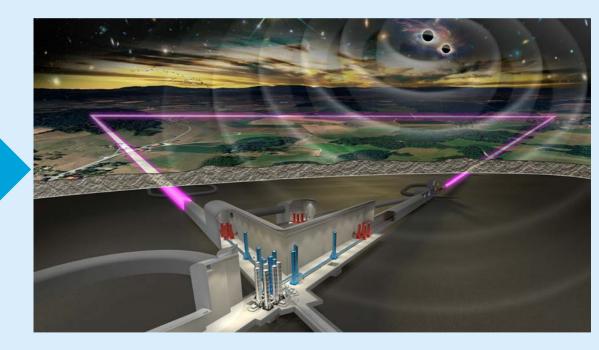


# Outline

- Overview of ETpathfinder
- Some examples of targeted R&D
- Current status and outlook



# From Current Detectors to 3G/Einstein Telescope



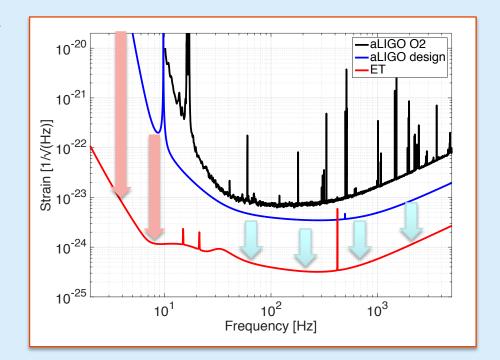
- ET can reward us with same of the most amazing Science!
- First though we have to solve many, many challenges!

# **Low Frequency Challenge**

At mid and high frequency we aim for factor ~10 improvement. At low frequency we are aiming for factors 100, 1000 and more improvement. -> Huge Challenge

Need to do fundamental changes in technology and concepts. => Need

technology and concepts. => Need to test, demonstrate, engineer and optimise those!



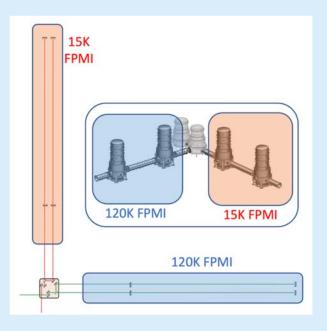
This is where we believe ETpathfinder can play a role!



### **Aim of ETpathfinder**

- Provide a low-phase noise, easy access, fast turnaround interferometer testbed for the Einstein Telescope.
- Allows tests at the system level (not just individual subsystems on their own) and to learn about dependency and interplay of different subsystems.
- Can facilitate an interferometer with 4 cryogenic test masses of 100+kg.
- Will run initially two independent interferometers with small mirrors (15cm diameter) to explore the ET/Voyager/CE matrix of cryogenic temperatures (123K and ~15K) and wavelengths (~2.1 and 1.55 microns).

**Maastricht University** 



# **New Technologies**

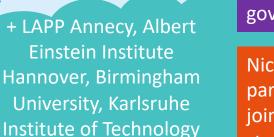


#### ET requires technological advances on all fronts:

- New mirror material => Silicon
- New temperature => 10-20K
- New laser wavelength => 1.5-2.1 microns
- Advanced quantum-noise-reduction schemes

# **ETpathfinder = Collaborative Effort!**

- 1. Nikhef
- 2. Maastricht University
- 3. Eindhoven University of Technology
- 4. University of Leuven
- 5. Ghent University
- 6. University of Antwerp
- 7. University of Hasselt
- 8. University of Liège
- 9. Vrije Universiteit Brussel
- 10. Université catholique de Louvain
- 11. Fraunhofer Institute for Laser Technology (ILT)
- 12. RWTH Aachen University
- 13. University of Twente
- 14. Flemish Institute for Technological Research (VITO), Mol
- 15. Netherlands Organisation for Applied Scientific Research (TNO), Delft



etc

€14.5m capital investment (Interreg, institutions, governments, provinces)

Nice side effect: Many new partners from outside GW joined (and some now also started to contribute to current detectors, i.e. Virgo.)

Open for everyone to join in the effort!

#### **German Universities funding to support ET + ETpathfinder**



#### The Federal Ministry of Education and Research supports the development of the Einstein Telescope

#### For the next generation of gravitational-wave detectors on Earth: Laser development and squeezed-light research in Hannover

#### JULY 24, 2020

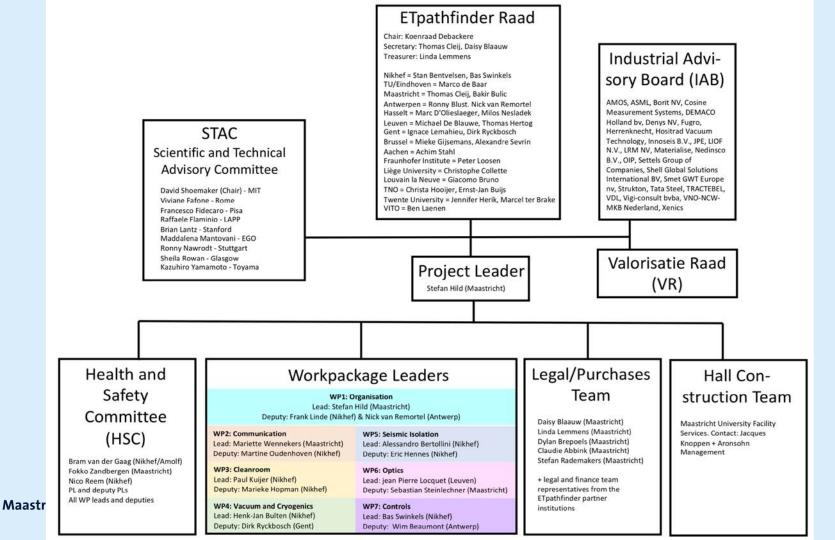
Fourteen German universities and non-university research institutions will receive a total of approximately 2 million Euros from the Federal Ministry of Education and Research (BMBF). This funding enables physicists and geophysicists to conduct research on various technological aspects of the Einstein telescope. These include, for example, seismic measurements and the development of crystalline fibers for suspending the mirrors.

At the Institute for Gravitational Physics of Leibniz Universität Hannover, in close cooperation with the Max Planck Institute for Gravitational Physics and the QuantumFrontiers Cluster of Excellence, a laser source for the Einstein Telescope Pathfinder – a prototype of the observatory – is to be developed and tested in Hannover and installed in the prototype in Maastricht. In addition, the participating researchers will investigate questions concerning the generation of squeezed light and the use of this technology in 3<sup>rd</sup> generation gravitational wave detectors. The BMBF funds will be used to finance an additional doctoral student position in Hannover.

Read more on about this in a news item on @ einsteintelescope.nl.

#### Spokespersons of the Verbund: A.Stahl (RWTH)





Slide 9

#### Workpackage Leaders

WP1: Organisation Lead: Stefan Hild (Maastricht) Deputy: Frank Linde (Nikhef) & Nick van Remortel (Antwerp)

WP2: Communication Lead: Mariette Wennekers (Maastricht) Deputy: Martine Oudenhoven (Nikhef)

WP3: Cleanroom Lead: Paul Kuijer (Nikhef) Deputy: Marieke Hopman (Nikhef)

WP4: Vacuum and Cryogenics Lead: Henk-Jan Bulten (Nikhef) Deputy: Dirk Ryckbosch (Gent) WP5: Seismic Isolation Lead: Alessandro Bertollini (Nikhef) Deputy: Eric Hennes (Nikhef)

WP6: Optics Lead: jean Pierre Locquet (Leuven) Deputy: Sebastian Steinlechner (Maastricht)

WP7: Controls Lead: Bas Swinkels (Nikhef) Deputy: Wim Beaumont (Antwerp)

#### STAC

#### Scientific and Technical Advisory Committee

David Shoemaker (Chair) - MIT Viviane Fafone - Rome Francesco Fidecaro - Pisa Raffaele Flaminio - LAPP Brian Lantz - Stanford Maddalena Mantovani - EGO Ronny Nawrodt - Stuttgart Sheila Rowan - Glasgow Kazuhiro Yamamoto - Toyama

#### From ETpathfinder Advisory Board (STAC) report

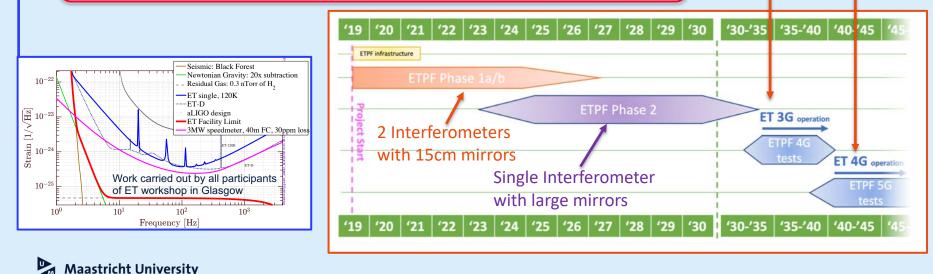
[...] Overall, the ETPF-STAC was very impressed with the vision for the facility, the technical capability of the leader and team, and the scope of the effort. It will be transformative for the field to have a facility and a research program covering the foreseen capabilities of the installation, and it can become a very natural center for technical innovation and scientific breakthroughs in precision measurement, interferometry, cryogeny for gravitational-wave detectors, and for the formation of a next generation of gravitational-wave scientists (to handle the next generation of gravitational-wave detectors). The growth of the team (and of the institutions interested in participating) is an exciting development and speaks to the timeliness and centrality of this infrastructure. [...]

The ETPF-STAC is very excited to be part of the establishment and exploitation of this unique facility and this dynamic team.



#### **ETpathfinder is a longterm activity!**

- ESFRI application states ET will be operational from 2035 to 2085.
- Expect many ET detector upgrades over the 50 years.
- While ET operates and observes in "generation X technology" ETpathfinder can do R&D for "generation X+1 technology"



# Outline

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- Some examples of targeted R&D
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#### ETpathfinder DESIGN REPORT

The ETpathfinder Team\*

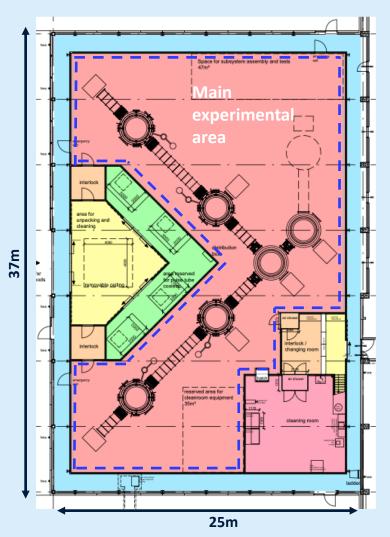
\*NEkhet, Maastricht University, University of Antwerp, Ghent University, Katholieke Universiteit Eusven, Université Catholique de Louvain, Hossett Iniversity, Vrije Universiteit Brussel, Frourhofer Institute for Loser Technology RWTH Aachen, University of Twente, Endhorven University of Technology. Liege Université, VITO, TNO.



In this talk I will only be able to present appetizers. Please see ETpathfinder for more details: <u>https://www.etpathfinder.eu/wp-</u> content/uploads/2020/03/ETpathfinder-Design-Report.pdf

### **Cleanroom Infrastructre**

- Inner dimensions of cleanroom (class 8, volume of 6000m<sup>3</sup>, 10 circulations/h) is 34.3m x 22m x 8.1m.
- Two separate foundations ---- (both 45cm thick concrete sitting on top of 170 piles drilled 5.8m into the ground) to decouple main experimental area from rest of building, AC, pumps etc.
- Special emphasis in tender on: 1) quiet air handling and filter units 2) improved acoustic isolation of cleanroom walls.
- Design includes a special compartment for noisy equipment (pumps, cryocoolers etc )

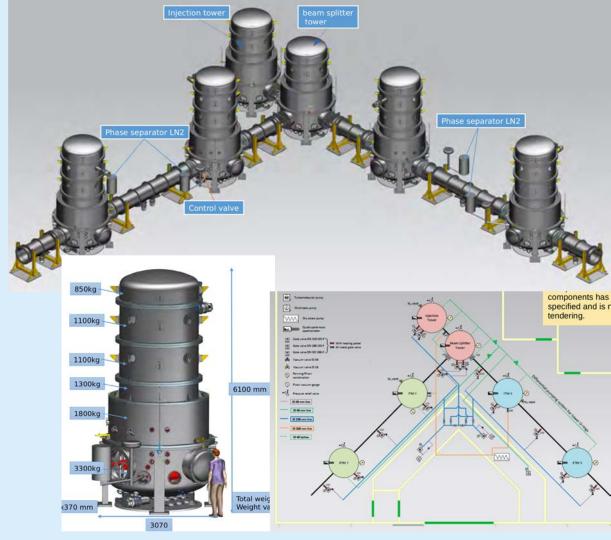


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### Vacuum

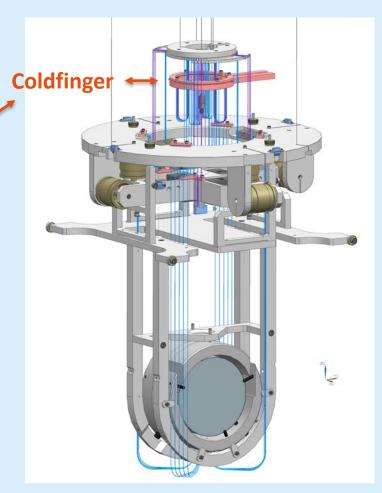
- 6 towers (4 mirror towers and 2 table towers); 80cm tubes; 20m arm length;
- Total volume about 130m<sup>3</sup>
- Target pressure of 10<sup>-9</sup>mbar
- Aiming for fast and easy access, i.e. to allow frequent venting etc: 1) access via mantle ring; 2) need to deal with lots of water => plan to have possibility to bake towers on regular basis at 90deg C.

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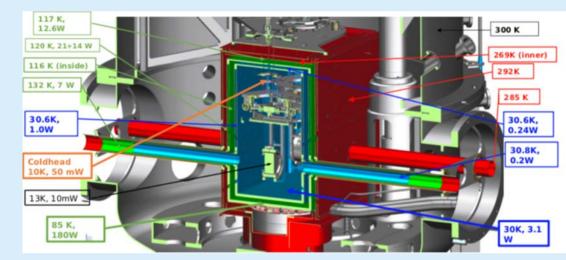
# **Cryogenics (I)**

- Mirrors need to be cooled to cryogenic temperatures (~15K, 123K), without introducing noise, i.e. cooling only possible via thin suspension wires.
- General approaches under consideration:
  - Dry system: pulse-tubes. Challenge = reduce and isolate vibrational noise.
  - Sorption coolers (base line in ETpathfinder) = more quite, less cooling power.
  - Cryogenic Liquids: LN2, He, He-II. Challenges
    avoid bubbling; transfer liquids from surface 300m above the caverns ...



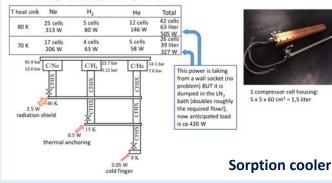
# **Cryogenics (II)**

 Need avoid ice on the mirrors and find ways to deal with ice if it builds up (reduce its thermal noise and optical influence)? – Will use 3 pairs of metal-cryo-shields.

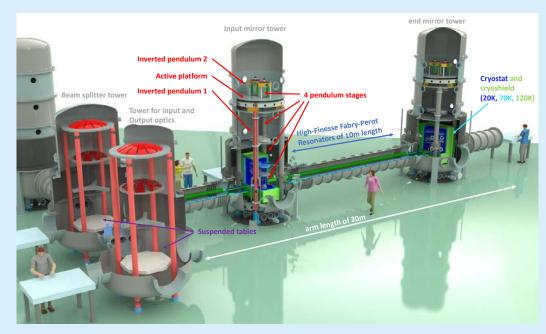


- Complex heat extraction matrix (vastly different powers, at different temperatures and with different noise requirements).
- Not only steady state operation sets requirements, but many come from cool-down requirements.
- No off-the-shelf simulation tools available that cover essential functionality. <sup>(3)</sup>

ELT design-based (3 stages, cold finger at 8K)

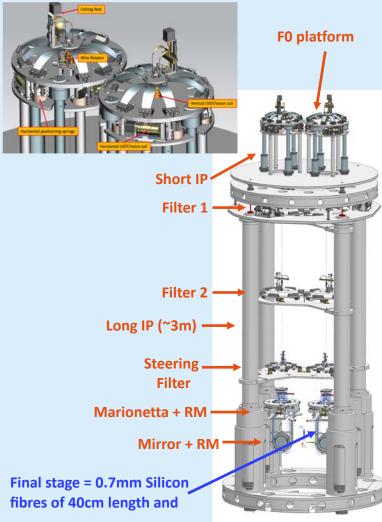


### **Seismic Isolation**

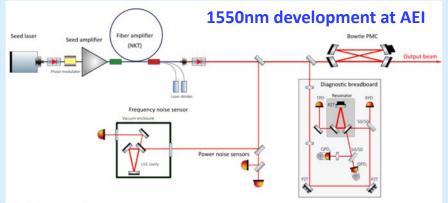


2 different designs: Mirror towers and table towers





### Lasers: 1550 nm and 2090nm



Setup for the measurement. AEI gitlab: fathie\Laser1550\PMC\PMCnoise\27102020\Setup

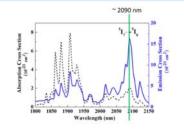


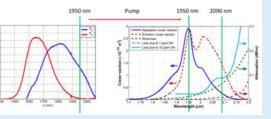
Planned setup (Fraunhofer ILT / E-TEST):

- Ring resonator based on Ho:YAG
- Single-frequency operation
- Emission wavelength 2090nm
- Output power 4W cw

Amplifier:

- Holmium fibre
- Pumped by Thulium fiber laser/amplifier



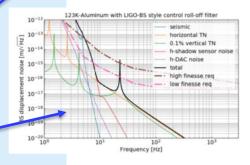


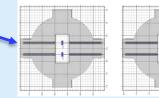


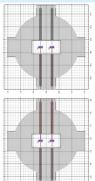
# **Optics + Coatings**

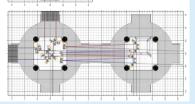
- Main mirrors (silicon 15cm diameter, 8cm thickness, FZ, R>14kOhm cm)
- First interferometer will explore currently available coatings (Silica/tantala? amorphous silicon? Multi-material coatings etc).
- Longer-term initiative @ KU-Leuven: Setting up an MBE system (crystalline coatings of different metal oxides)
- Arm cavity Finesse chosen as 800 (trade-off between lock-acquisition, compatibility of QN reduction schemes, Seismic isolation requirements of CITF etc.)
- Optical layout for 2 interferometers (one per arm)













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| Code               |   | Title  |                                    | Date         | Author(s)   |
| ET-00478-20        | CITY M  | Mechanical Stability of Vacuum Body                |                                    | 16/11/20     | M. Baars  |
| ET-0054A-20        | E117  | ETPF electronics tasks, weekly meeting 2020-08-24  |                                    | 02/11/20     | 8. Swinkels   |
| ET-0053A-20        | CTIV ET                                       | ETPF presentation for Nikhef Computing, 2020-07-07 |                                    | 02/11/20     | B. Swinkels   |
| ET-0052A-20        |   | Pathfinder thoughts al<br>19-06-17                 | bout controls, weekly meeting      | 02/11/20     | B. Swinkels   |
| ET-0051A-20        |   | Pathfinder in Maastric<br>19-05-21                 | ht, presentation at GWADW Elba,    | 02/11/20     | B. Swinkels   |
| ET-0050A-20        |   | esentations from first p<br>astricht, 2018-05-09   | prototype brainstorm meeting in    | 02/11/20     | J. v.d. Brand, B. Swinkels, H<br>Lück, S. Hild, C. Collette |
| ET-0049A-20        |   | Management Organisa<br>x.eu/tds/ql/3c=15703)       | tion (disabled, see https://apps.e | et- 02/11/20 | Michele Punturo on behalf of<br>the Governance Committee    |
| ET-0048A-20        | ETT ET  | PF_CR-2020-02 Vacuum                               | n system changes Feb-Oct 2020      | 22/10/20     | E. Hennes, M. Baars, K. Lam                                 |
| ET-00448-20        | 677 20  | 20-05-14 Galling preve                             | ntion measures                     | 22/10/20     | M. Baars  |
| ET-0046A-20        |   | 20-07-08 Gailing preve<br>-00448)                  | ntion measures 3 (included in      | 21/10/20     | M. Baars  |
| ET-0045A-20        | er 20<br>env Et                               | 20-05-20 Gailing preve<br>-00448) MITTICTUD ACCID  | ntion measures 2 (included in      | 21/10/20     | M. Baars  |
| ET-0043A-20        | 17<br>20                                      | 20-04-08 Baking and ts                             | slation model                      | 21/10/20     | M. Baars  |
| ET-0042A-20        | CTW 20  | 20-03-19 Design change                             | es from FEA                        | 21/10/20     | M. Baars  |
| ET-0041A-20        |   | ange Request ETPF-CR<br>recussion Compensation     | 2020-01-Large IP - Inverted        | 20/10/20     | E. Hennes, K. Lam, M. Baars                                 |
| ET-0040A-20        | ETHY ANA<br>ETHY ANA<br>ETHY ANA<br>ETHY ITAC | ides for ETpathfinder c                            | uli 2020-10-19                     | 19/10/20     | Stefan Hild   |
| ET-0039A-20        | CTW E   | PathFinder Benchtop S<br>PFworkshop 8-9/10/20      | uspensions Update (for<br>20)      | 12/10/20     | Joris van Heljningen  |
| ET-0038A-20        | er<br>ciw                                     | pathfinder payload set                             | smic residual vibration            | 10/10/20     | E. Hennes   |
| ET-0037A-20        | er<br>env                                     | pathfinder optical laye                            | ut update                          | 09/10/20     | S. Steiniechner   |
| ET-0036A-20        | er<br>env El                                  | pathfinder noise budg                              | et update                          | 09/10/20     | Andrei Utina  |
| ET-0035A-20        | 17  | ol-down curves for ETp                             | athfinder                          | 08/10/20     | H.J. Bulten   |

Again in this talk I will only be able to give appetizers of current activities. If interested please follow activities via ET-TDS: <u>https://apps.et-gw.eu/tds/</u>



#### **Timelaps of hall re-construction**



https://youtu.be/gtaJ5P16mCw





#### **Status: Cleanroom**





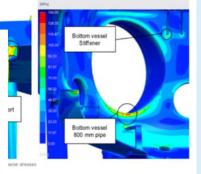
#### Vacuum system tender

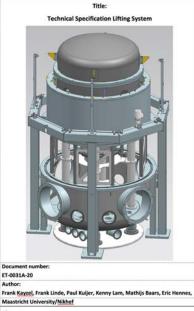
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| Code             |                   | Title  |                               | Date     | Author(s)   |
| T-0070A-20       | ET Te             | chnical construction fil                             | e vacuum vessel tender        | 24/11/20 | M. Baars for the ETpathfinder<br>team   |
| T-0047B-20       | ET MG             | chanical Stability of Va                             | acuum Body                    | 16/11/20 | M. Baars  |
| T-0048A-20       | ETPF              | PF_CR-2020-02 Vacuum                                 | system changes Feb-Oct 2020   | 22/10/20 | E. Hennes, M. Baars, K. Lam   |
| T-0044B-20       | ET<br>ETPF 202    | 20-05-14 Galling prever                              | ntion measures                | 22/10/20 | M. Baars  |
| T-0046A-20       |                   | 20-07-08 Galling prever<br>-0044B) RESTRICTED ACCESS | ntion measures 3 (Included in | 21/10/20 | M. Baars  |
| T-0045A-20       |                   | 20-05-20 Galling prever<br>-0044B) RESTRICTED ACCESS | ntion measures 2 (Included in | 21/10/20 | M. Baars  |
| T-0043A-20       | ET 20             | 20-04-08 Baking and isc                              | olation model                 | 21/10/20 | M. Baars  |
| T-0042A-20       | ET<br>ETPF 20     | 20-03-19 Design change                               | rs from FEA                   | 21/10/20 | M. Baars  |
| Г-0041А-20       |                   | ange Request ETPF-CR-<br>rcussion Compensation       | 2020-01-Large IP - Inverted   | 20/10/20 | E. Hennes, K. Lam, M. Baars   |
| Г-0031А-20       | ET Te             | chnical Specification Li                             | ifting System ETpathfinder    | 27/08/20 | Frank Kayzel, Frank Linde,<br>Paul Kuijer, Kenny Lam,<br>Mathijs Baars, Eric Hennes                           |
| Г-0030A-20       | ст Те             | chnical Specification V                              | acuum Vessel ETpathfinder     | 27/08/20 | Frank Kayzel, Frank Linde,<br>Paul Kuijer, Stefan Hild,<br>Kenny Lam, Mathijs Baars,<br>Henk-Jan Bulten, Eric |

Maastricht University

The shown stresses (Figure 14<sup>6</sup>) vary per case ('a'/b') and with different mesh methods and sizes and with different connection methods, but are always larger than the allowed 140 MPa. The only exception is the vessel support, Inter-ring which does not give excessive Stiffener stresses when glue is used instead of modeling the bolt. The bottom vessel stiffener only shows excessive stress in case 'b'. The 800 mm pipe weld has shown stress for both cases7. The elaborate review of the results is given in Appendix 2 - with corresponding screenshots shown

in Appendix 6 -





Slide 27

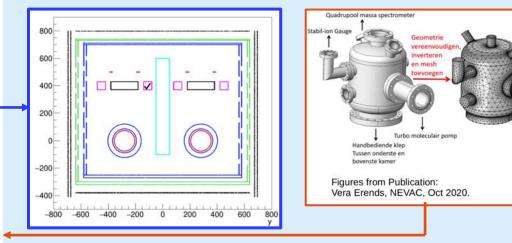
Abstract

The object of this Technical Specification is to describe the requirements of the Lifting System for

# Vacuum / cryogenic simulations

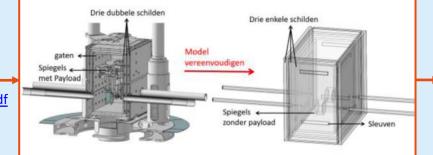
#### Modeling:

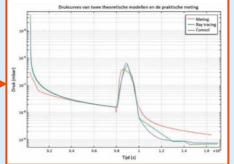
- ETpathfinder is a test facility: pump-down should be quick and part of the system will be vented quite often. Largest challenge: water (monolayer of water that will bind to the surface after each venting).
- In CDR: we developed a simulation package that tracks water molecules : adsorption and desorption on the walls, permeation through Viton O-rings, molecular flow, compression factor of turbo pumps etc.
- Molflow and Comsol were not capable of doing these calculations (no pressure-dependent, time-dependent and coverage-dependent parameters for the surfaces, and the ETpathfinder geometry was too detailed).
- Calibrated the calculations against Molflow and Comso using a outgassing set-up at Nikhef (Vera Erends, Berend Munneke, HJB)



Pics from Article by Vera: <u>https://nevac.nl/archief\_pdf/pdf\_208.pdf</u>

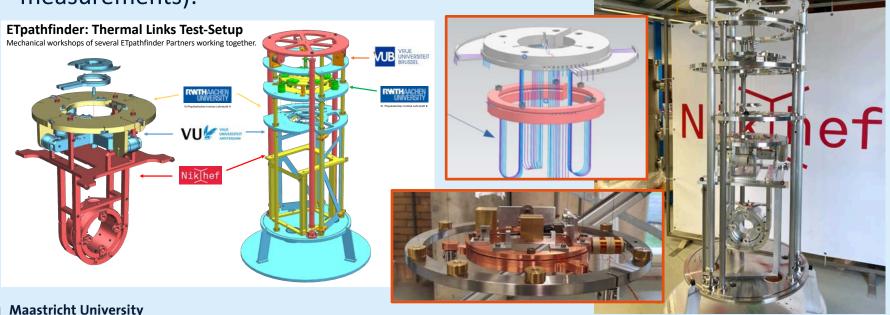






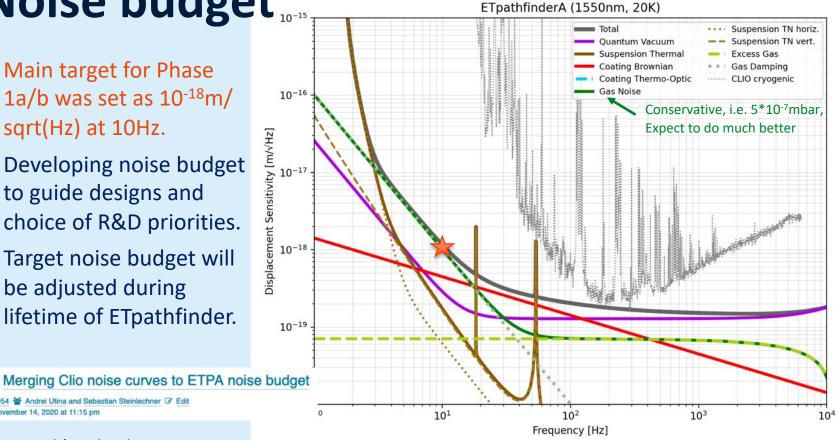
#### **Test setups**

- Various test setups are underway for cryogenics and seismic isolation.
- A nice example of collaboration combining efforts from several mechanical workshops (Jelly fish heat link mechanical coupling measurements):



# Noise budget

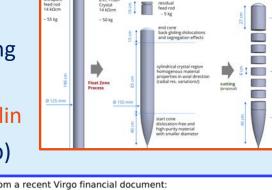
- Main target for Phase 1a/b was set as 10<sup>-18</sup>m/ sqrt(Hz) at 10Hz.
- Developing noise budget to guide designs and choice of R&D priorities.
- Target noise budget will be adjusted during lifetime of ETpathfinder.





### Many things on the go ...

- Tenders in preparation for: 1) large bellows, 2) vacuum pumps and diagnostics tools, 3) Gate valves, 4) MBE coating system, 5) Optical tables ...
- Silicon mirror substrates currently in production at IKZ Berlin
- Vacuum system lifting system final design checks (Antwerp)
- Control system assembly at LAPP
- 1550nm laser being set up at AEI
- 2090nm laser source under development at Fraunhofer Institute Aachen
- Silicon polishing tests underway at Vrije University of Brussels



LEIBNIZ-INSTITUT FÜR KRISTALLZÜCHTUNG

2×

Benefits go both ways, from a recent Virgo financial document: "DAQ.03: Mezzanine ADC2378 production: The production is cheaper than expected. The production of ADC mezzanines for AdV+ was carried out at the same time as the production of ETpathfinder. A higher volume of cards lowers the unit price of a card and lowers the price for AdV+."

2x



### **Take-away-points**

- ETpathfinder is reality now (you can see, touch, breath it, if you stop by in Maastricht).
- ETpathfinder is a long-term initiative and independent of the site decision (i.e. ETpathfinder will not vanish in 2025, but be available to the ET community for the next decades.)
- ETpathfinder is huge opportunity to test new technology vital for ET.
  Let's use it to increase (our and the communities and funder's) confidence into ET.
- Everybody is welcome to join the ETpathfinder team. Lots of work available for new groups to take charge of. Please get in contact.



# Thank you for your attention.

#### Any questions?

(Please ask, but also feel free to get into contact via email at any time!)





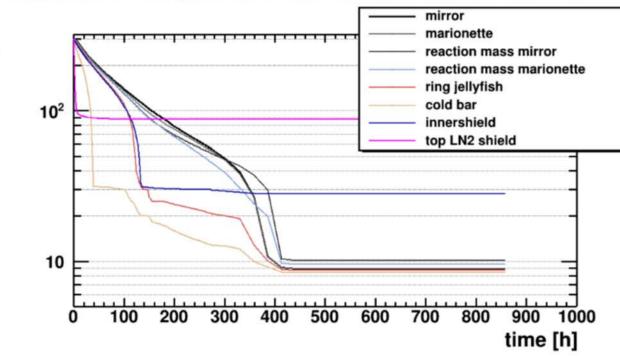
# **EXTRA SLIDES**



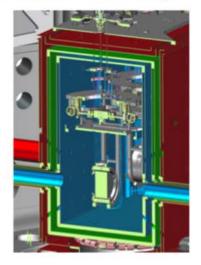
#### Cool-down curves

Calculations with 30W cooling power (both shields and cold finger) above 30K

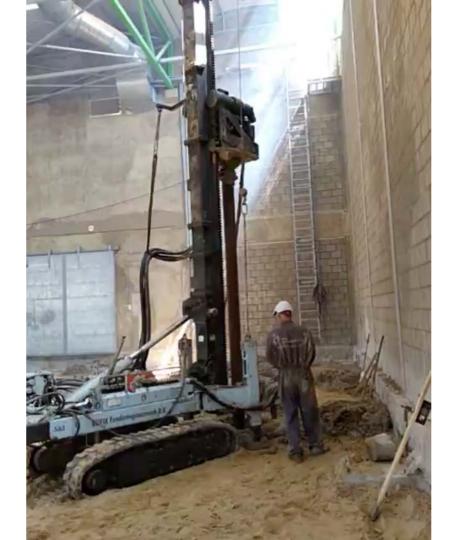
Temperature as a function of time for several elements. Results for high-emissivity coatings. Final temperatures and equilibrium are reached in 400 h, about 5 times quicker than with low-emissivity

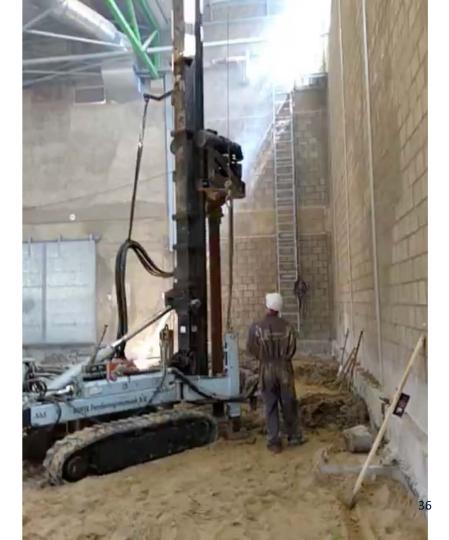


temperature [K]



Masses mirror 3.3kg, marionette 8kg, reaction masses 10 and 5 kg, cold finger 20 kg, 70kg of cryogenic mass. Shield weigh about 130 and 150 kg







Lots of space for LN2 tanks outside the hall (with underground Tunnel for pipes into the hall)

#### New 360kW cooling water unit installed on the roof of neighboruring hall

Carrier

DAIKIN

11 11



#### From STAC report, Feb 2020

The interactions with the other major efforts in the GW field are important for the ETPF. The Einstein Telescope Project has a lot of overlap with staff and objectives; it would be good to have a compact description of that special relationship (is ETPF obligated to pursue priorities determined by ET leadership? Is the ET Project obligated to adopt designs developed at ETPF?). It would be advantageous that the ETPF also be seen by the US Cosmic Explorer and Voyager undertakings as a partner to address technical issues and to exchange staff and visitors, and we encourage ETPF to invest some energy in being visible to that community. Further steps could be to organize a joint meeting at Maastricht of LSC-Virgo-KAGRA scientists engaged in 3G detectors, and/or to suggest for a US LVK meeting to have a significant session on ETPF where input to the program is invited (and the program of course presented).

The 'joint citizenship' in LSC and Virgo is a good first step. The plan for Virgo is to have those interested in ET to have a significant engagement in today's Virgo; this feels like it is compatible with the ETPF planning. It is important to continue the discussions in both the LSC and Virgo on 'credit' for 3rd generation detectors and the relationships that enforces between 3G detectors and LSC-Virgo. KAGRA has unique overlap with the ETPF scope in addressing the challenges of ~20K cryogenics for GW detectors, and a deep relationship between KAGRA and ETPF would appear to be mutually rewarding.

The ETPF-STAC recommends that there be a written document describing the relationship between ETPF and the ET project (this might end up into an MoA at some point).

